



Influence of Prior Covariance Structure on Inverse Estimates of CO₂ Fluxes in Los Angeles Basin

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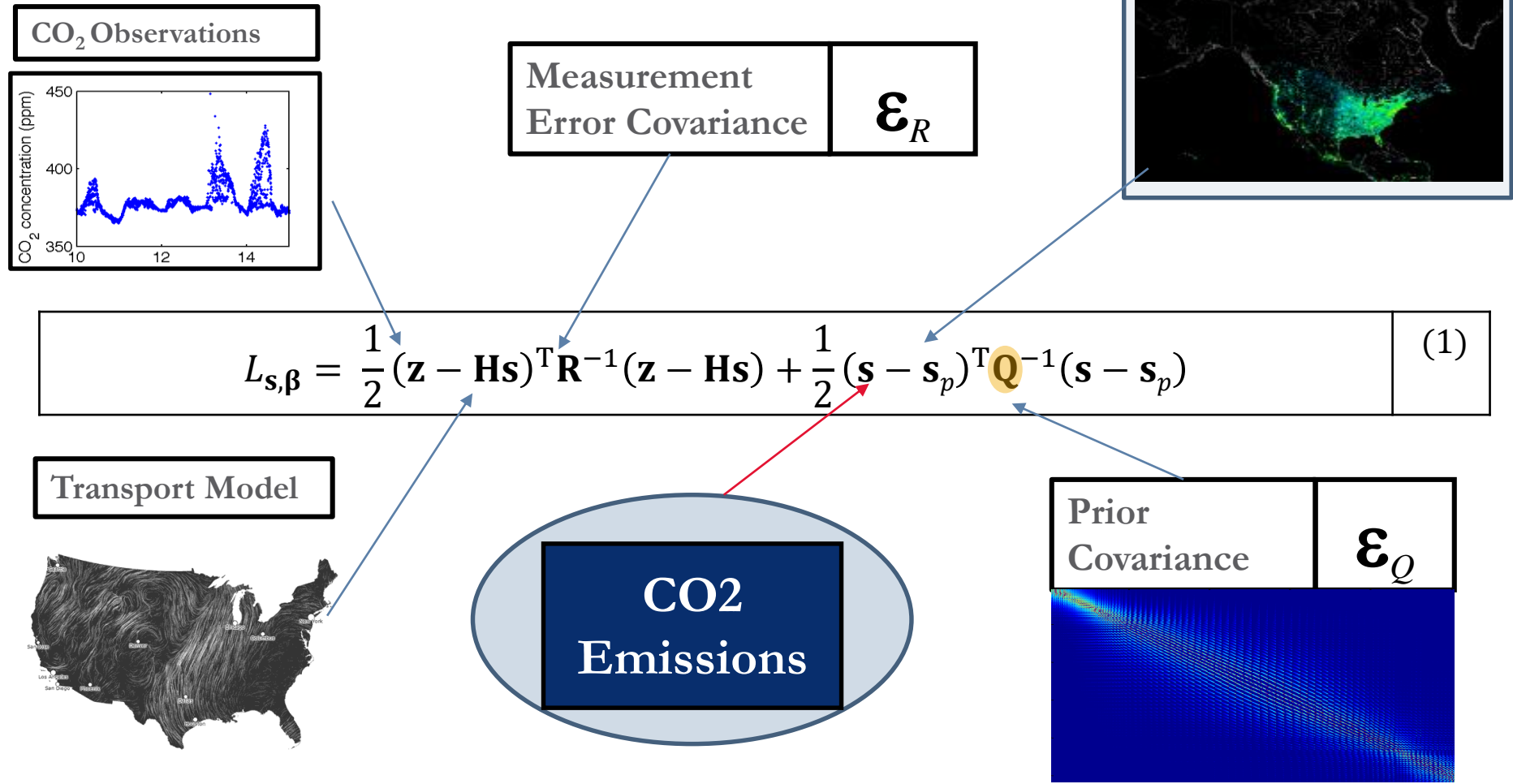
Jet Propulsion Laboratory
California Institute of Technology

Layout of the presentation

- Formulation of atmospheric Inverse Models
- Criteria for assessing inverse models
- Choices that impact inverse models
- Types of prior covariance used in inverse models
- The phenomenon for which prior covariance needs to be defined
- Role of prior covariance in inverse output
- Case Studies:
 - Regional: North America
 - Urban: Los Angeles

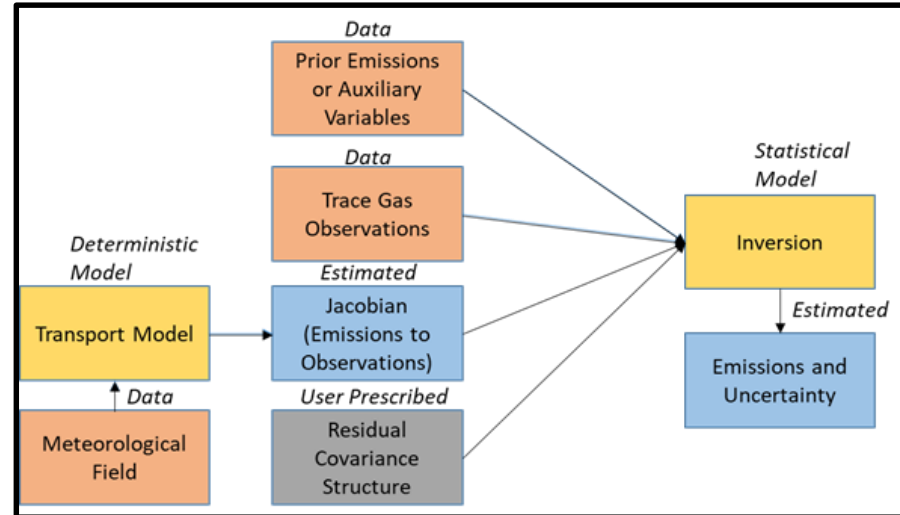
Statistical Approach to Atmospheric Inversions

Atmospheric Inversions: Components of (linear) Statistical Model



Flavors of Atmospheric (linear) Inverse Models

Inverse Process



Bayesian

$$L_s = \frac{1}{2}(\mathbf{z} - \mathbf{H}\mathbf{s})^T \mathbf{R}^{-1}(\mathbf{z} - \mathbf{H}\mathbf{s}) + \frac{1}{2}(\mathbf{s} - \mathbf{s}_p)^T \mathbf{Q}^{-1}(\mathbf{s} - \mathbf{s}_p) \quad (2)$$

Geostatistical

$$L_{s,\beta} = \frac{1}{2}(\mathbf{z} - \mathbf{H}\mathbf{s})^T \mathbf{R}^{-1}(\mathbf{z} - \mathbf{H}\mathbf{s}) + \frac{1}{2}(\mathbf{s} - \mathbf{X}\boldsymbol{\beta})^T \mathbf{Q}^{-1}(\mathbf{s} - \mathbf{X}\boldsymbol{\beta}) \quad (3)$$

Another Formulation

$$L_{s,\beta,u} = (\mathbf{z} - \mathbf{H}\mathbf{s})^T \mathbf{R}^{-1}(\mathbf{z} - \mathbf{H}\mathbf{s}) + (\mathbf{s} - \mathbf{X}\boldsymbol{\beta} - \mathbf{M}\mathbf{u})^T \mathbf{Q}^{-1}(\mathbf{s} - \mathbf{X}\boldsymbol{\beta} - \mathbf{M}\mathbf{u}) + \mathbf{u}^T \mathbf{P}^{-1} \mathbf{u} \quad (4)$$

Criteria for assessing an inverse model (other than Uncertainty Reduction)

Correlation Coefficient and RMSE

$\text{corr}(\mathbf{z}, \mathbf{H}\hat{\mathbf{s}})$	(5)
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$\text{RMSE} = \sqrt{\frac{\mathbf{1}^T (\mathbf{z} - \mathbf{H}\hat{\mathbf{s}})^{\circ 2}}{n}}$	(6)
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Hat Matrix and Cross Validation

$\mathbf{h} = \mathbf{H}\mathbf{s}_p \left((\mathbf{H}\mathbf{s}_p)^T (\mathbf{H}\mathbf{Q}\mathbf{H}^T + \mathbf{R})^{-1} \mathbf{H}\mathbf{s}_p \right)^{-1} (\mathbf{H}\mathbf{s}_p)^T (\mathbf{H}\mathbf{Q}\mathbf{H}^T + \mathbf{R})^{-1}$	(7)
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$cv = \frac{1}{n} \sum_{i=1}^N \left(\frac{e_i}{1 - h_{ii}} \right)^2$	(8)
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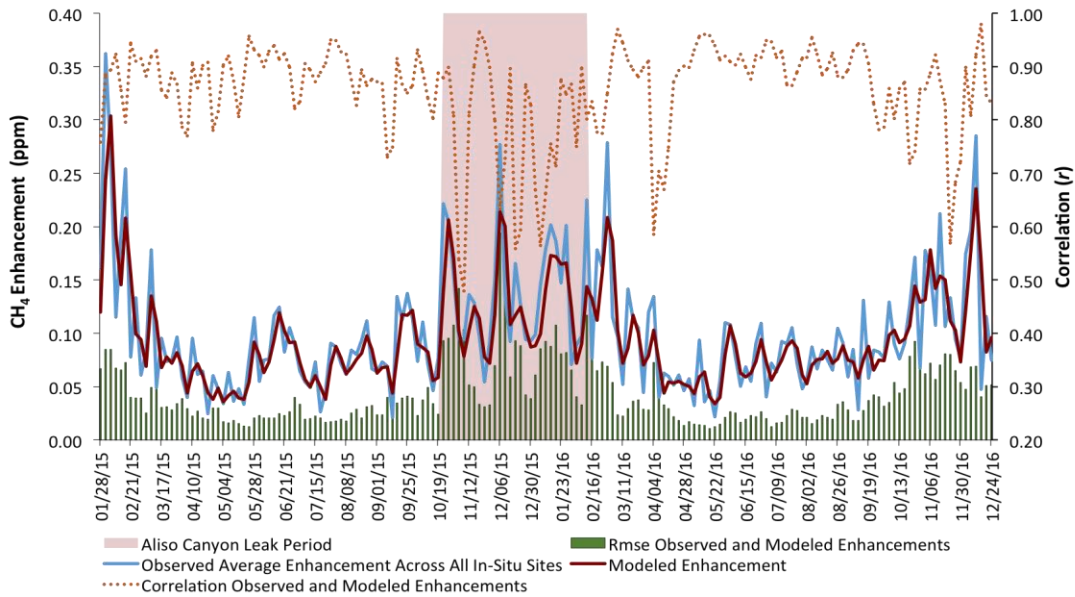
Averaging Kernel Matrix

$\mathbf{G} = \mathbf{Q}\mathbf{H}^T (\mathbf{H}\mathbf{Q}\mathbf{H}^T + \mathbf{R})^{-1} \mathbf{H}$	(9)
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Reduced Chi-Square Statistic

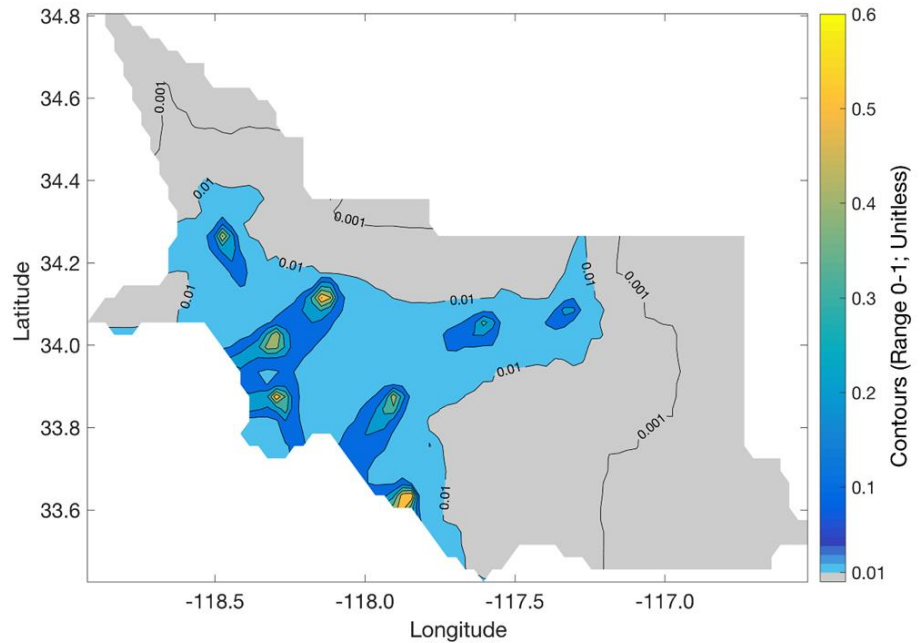
$\chi_{red}^2 = \frac{(\mathbf{z} - \mathbf{H}\hat{\mathbf{s}})^T \mathbf{R}^{-1} (\mathbf{z} - \mathbf{H}\hat{\mathbf{s}}) + (\mathbf{s} - \mathbf{s}_p)^T \mathbf{Q}^{-1} (\mathbf{s} - \mathbf{s}_p)}{n}$	(10)
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Criteria for assessing an inverse model: Examples



Correlation and RMSE

Averaging Kernel



Sensitivity Analysis

- How to determine which factor played most important role in influencing estimates of fluxes ?
- There are multiple ways to do this but partial derivatives provide a complete framework to do this.

$$\Psi = (\mathbf{H}\mathbf{Q}\mathbf{H}^T + \mathbf{R})$$

$$\frac{\partial \hat{\mathbf{s}}}{\partial \mathbf{z}} = \underbrace{\mathbf{Q}\mathbf{H}^T\Psi^{-1}}_{\text{Kalman Gain}} \quad (11)$$

$$\frac{\partial \hat{\mathbf{s}}}{\partial \mathbf{Q}} = \mathbf{H}^T\Psi^{-1}(\mathbf{z} - \mathbf{H}\mathbf{s}_{prior})\otimes \mathbf{I}_k - \mathbf{H}^T\Psi^{-1}(\mathbf{z} - \mathbf{H}\mathbf{s}_{prior})\otimes \mathbf{H}^T\Psi^{-1}\mathbf{H}\mathbf{Q} \quad (12)$$

$$\frac{\partial \hat{\mathbf{s}}}{\partial \mathbf{R}} = \Psi^{-1}(\mathbf{z} - \mathbf{H}\mathbf{s}_{prior})\otimes \Psi^{-1}\mathbf{H}\mathbf{Q} \quad (13)$$

$$\overset{\text{Normalized}}{\text{Sensitivity}} \Delta \hat{\mathbf{s}} = \frac{\kappa_i^0}{\hat{\mathbf{s}}(\kappa^0)} \times \left[\frac{\partial \hat{\mathbf{s}}}{\partial \kappa_i^0} \right] \quad (14)$$

Importance of Prior Covariance Matrix

Impact of Prior Covariance (North America Example)

Separable Exponential Space-Time

$$\mathbf{Q} = \sigma^2 \left[\exp \left(\frac{-\mathbf{d}_{temporal}}{l_{temporal}} \right) \otimes \exp \left(\frac{-\mathbf{d}_{spatial}}{l_{spatial}} \right) \right] \quad (10)$$

Spatially dependent error variance

$$\mathbf{Q} = \left(a \begin{bmatrix} k_1 & 0 & 0 \\ 0 & . & 0 \\ 0 & 0 & k_r \end{bmatrix} + b \begin{bmatrix} 1 & 0 & 0 \\ 0 & . & 0 \\ 0 & 0 & 1 \end{bmatrix} \right) \quad (11)$$

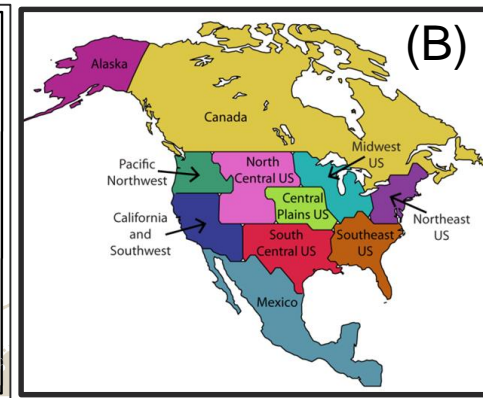
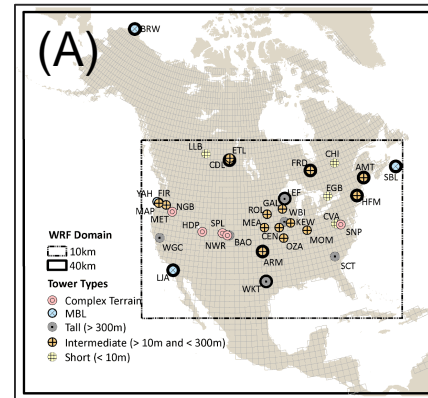
Assessment: BIC

$$BIC = \underbrace{RSS + \ln |(\mathbf{H}\mathbf{Q}\mathbf{H}^T + \mathbf{R})^{-1}|}_{\log \text{ likelihood}} + \underbrace{p \ln (n)}_{\text{penalty term}} \quad (12)$$

Impact of Prior Covariance (North America Example II)

Details of the Case Study:

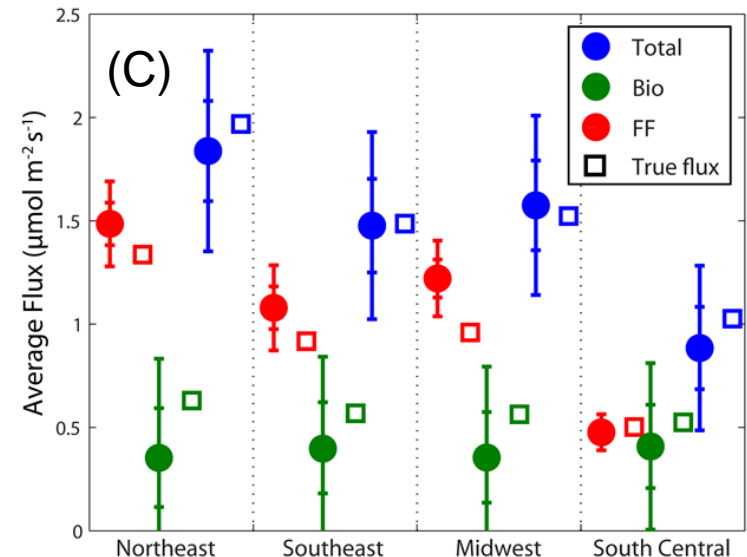
- Inversion Area: North America
- Inversion Time Period: 2008
- Resolution: 3-Hourly, $1^\circ \times 1^\circ$
- Observations: 35 in-situ towers
- Simulation Study: True Fluxes were known
- Prior Covariance Assessed:
 1. Night Lights
 2. Population Density
 3. Urban Area
 4. FF Inventory
 5. Separable Exponential Covariance



A. Study Area and In-Situ Towers
B. Flux Aggregation Area
C. Results from the case study

Results:

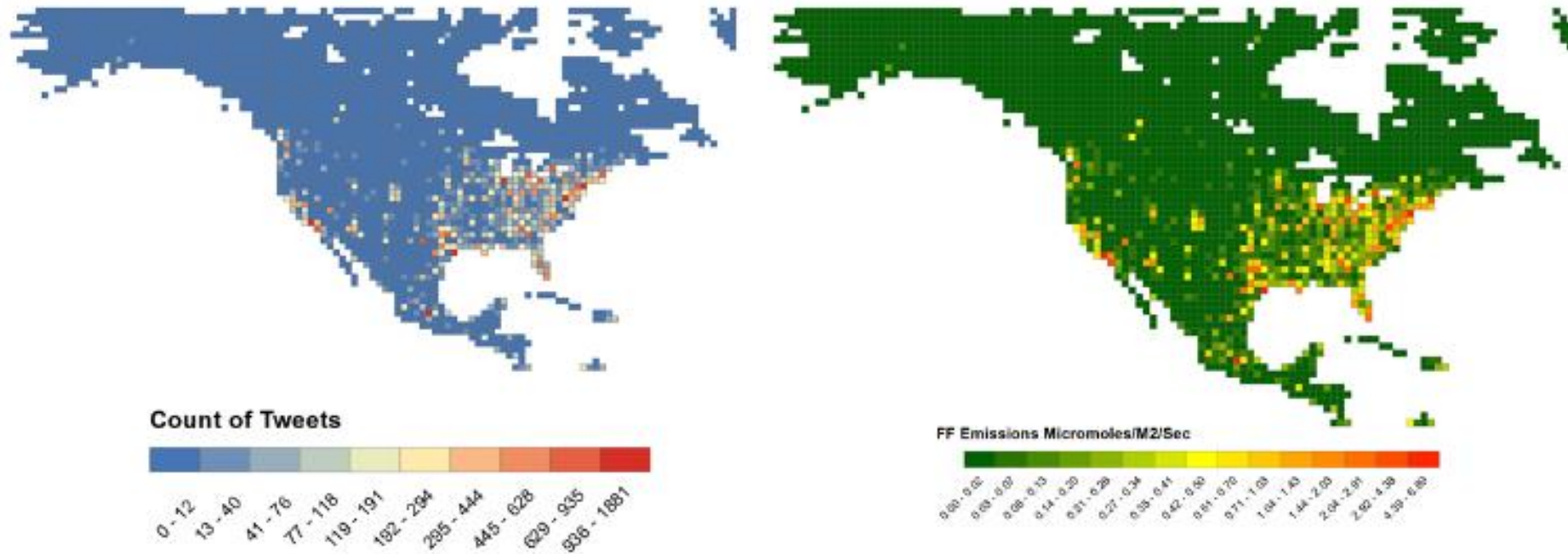
- FF Inventory based covariance considerably better than other covariance structures



Yadav et. al. (JGR-Atmospheres 2016)

Impact of Prior Covariance (North America Example III)

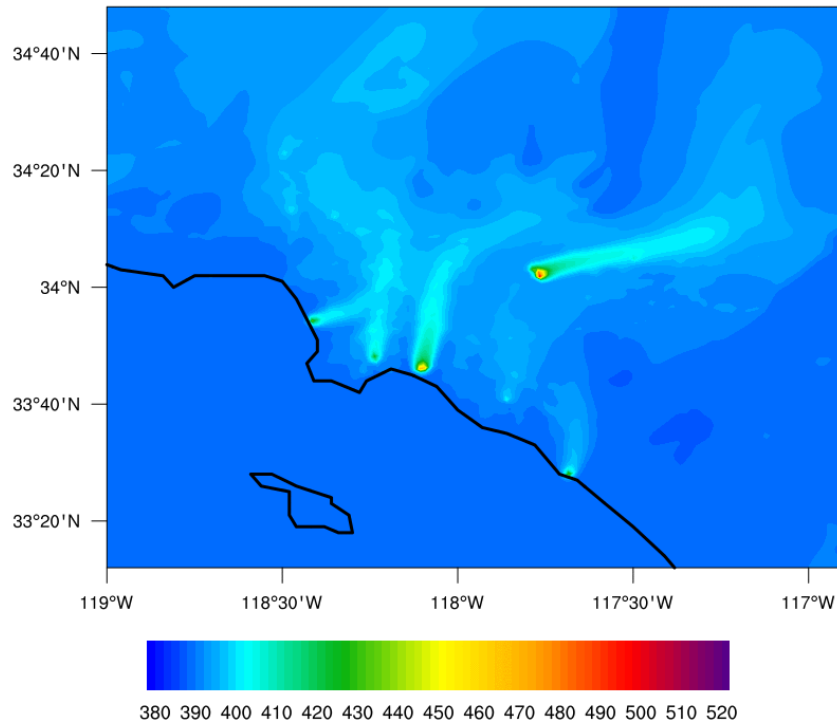
Work in Progress: Estimating Fossil Fuel Emissions By Using Twitter Feeds



Observations to Fluxes: Why is prior covariance so important in urban areas (Example from Los Angeles)

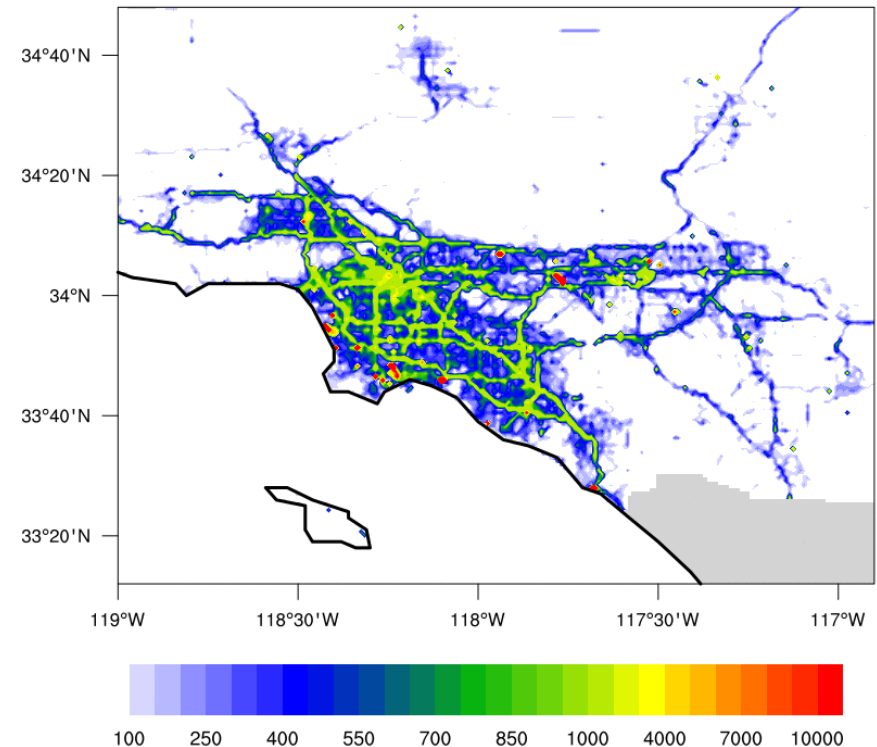
Atmospheric CO₂ Concentrations (parts per million)

2011-09-11 16:00:00 PST



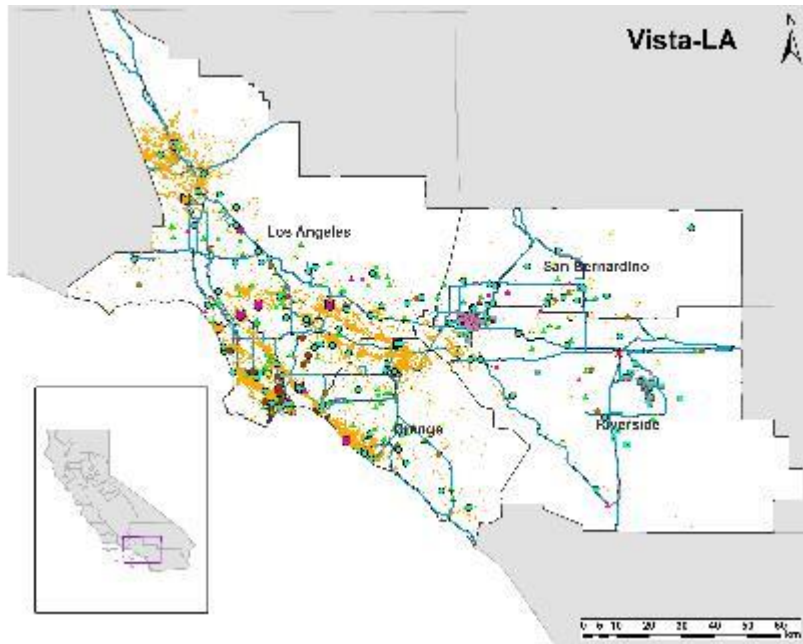
Fossil Fuel CO₂ Emissions (kilograms/hour)

2011-09-11 16:00:00 PST

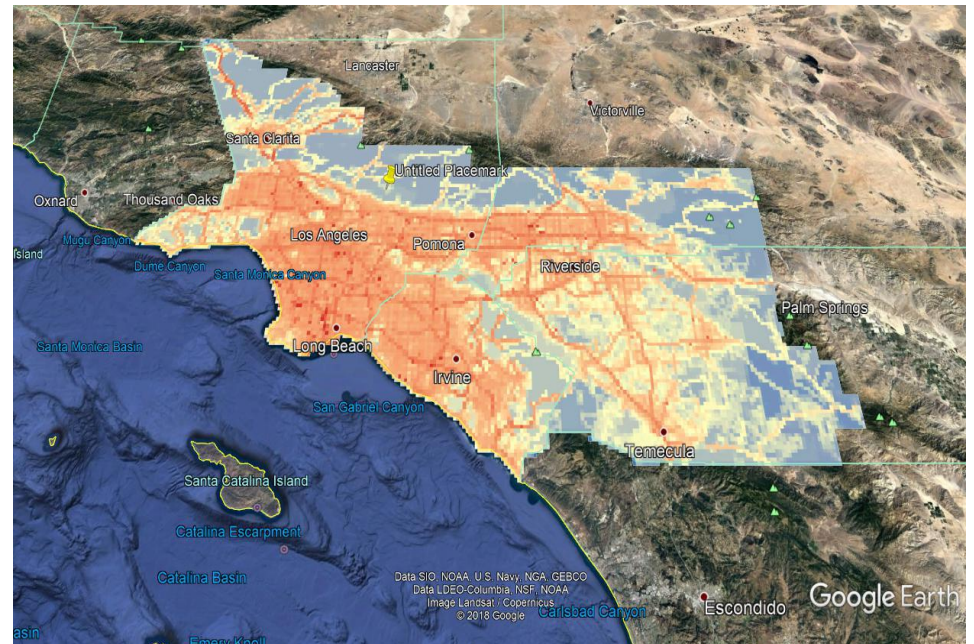


Observations to Fluxes: Why is prior covariance so important in urban areas (Example from Los Angeles) II

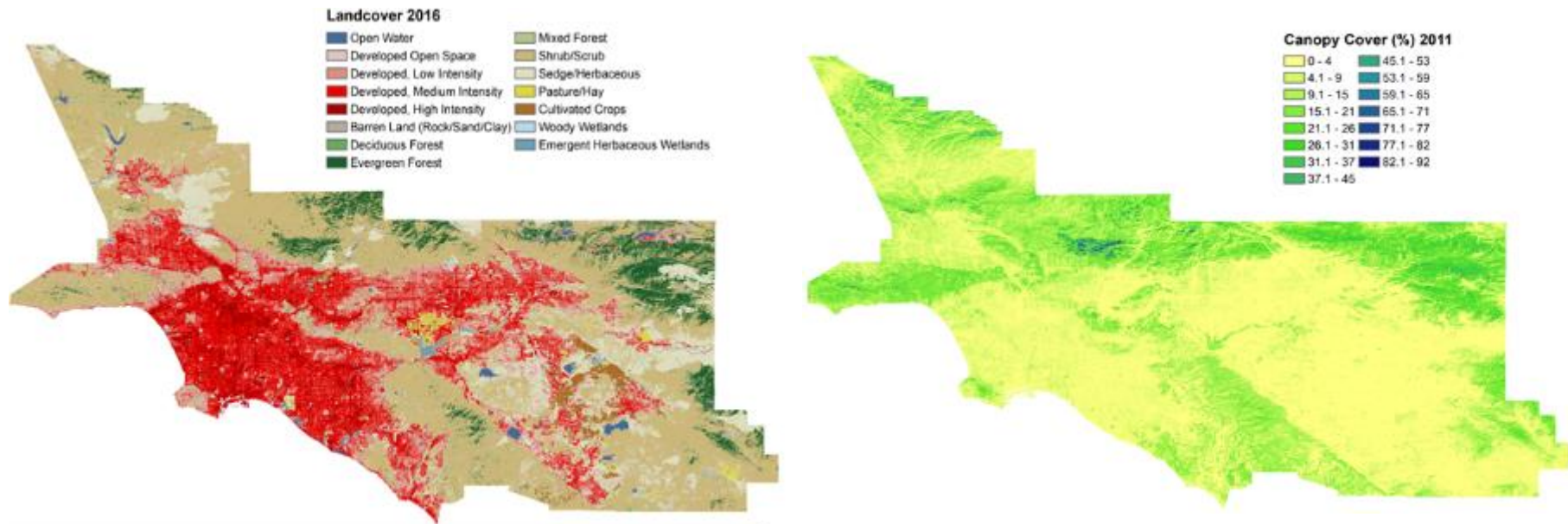
Distribution of Methane Emitting Infrastructure



Hourly Carbon dioxide emissions from Los Angeles



Observations to Fluxes: Why is prior covariance so important in urban areas (Example from Los Angeles) III



Behavior of different covariance formulations

Details of the Case Study:

- Inversion Area: Los Angeles
- Inversion Time Period: 2015
- Resolution: 4-day, 3km
- Observations: 6 in-situ towers
- Real data Study:
- Prior Covariance Assessed:
 1. FF Inventory (diagonal)
 2. Separable Exponential Covariance
 3. Temporal correlation and diagonal spatial

Results

- Correlation length in Space is non present
- Correlation length is considerably larger in time

Conclusions and Future Steps

- Implement proposed covariance structures for estimating fluxes

$\mathbf{Q} = \sigma^2 \left[\exp \left(\frac{-\mathbf{d}_{temporal}}{l_{temporal}} \right) \otimes \left(a \begin{bmatrix} k_1 & 0 & 0 \\ 0 & \cdot & 0 \\ 0 & 0 & k_r \end{bmatrix} + b \begin{bmatrix} 1 & 0 & 0 \\ 0 & \cdot & 0 \\ 0 & 0 & 1 \end{bmatrix} \right) \right]$	(13)
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- Include observations from multiple instruments



- Perform sensitivity analysis
- Use real time social media to better define temporal covariance model